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Ultraviolet Absorption Coefficients of CO2, CO,

 $0_2$ ,  $H_20$ ,  $N_20$ ,  $NH_3$ , NO,  $SO_2$ , and  $CH_4$  Between

1850 and 4000 A

Code 2A

by

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and

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(NASA CR-52089)

# UNPUBLISHED PRELIMINARY DATA

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Ultraviolet Absorption Coefficients of  $CO_2$ ,  $CO_3$ ,  $CO_4$ 

### Abstract

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The ultraviolet absorption coefficients have been determined for  $CO_2$ ,  $CO_3$ ,  $CO_4$ ,  $CO_4$ ,  $CO_5$ ,  $CO_4$ ,  $CO_5$ ,  $CO_$ 

### Introduction

Ultraviolet absorption coefficients reported in the literature fall, in general, into two categoriess those determined at wavelengths below 1850 A and those determined at wavelengths above 2500 A. The intermediate region, between 1850 and 2500 A is of great interest in connection with photochemical studies of planetary atmospheres since the high intensities of the solar radiation and the long path lengths through the atmospheres may result in significant amounts of photochemical reactions even where the absorption coefficient is very small. The excellent work of Watanabe, Zelikoff, and Inn (1953) covers this wavelength range for many gases, but, in general, does not give values below 0.1 cm<sup>-1</sup>. For these reasons the absorption coefficients of a number of gases which are known to be either major or minor constituents of various planetary atmospheres were determined

### **Experimental**

Description of Apparatus: Absorption experiments were carried out using a Perkin-Elmer Model 350 absorption spectrophotometer. This is a double beam instrument covering the spectral region from about 1850 A to 2.7  $\mu$ . For these studies measurements were made between 1850 and 4000 A. Using a 10 cm path length and a scale expansion of 50 it was possible to measure absorption coefficients as low as  $10^{-4}$  cm<sup>-1</sup>.

<u>Purification of Gases</u>: The gases studied were  $CO_2$ ,  $CO_3$ ,  $O_2$ ,  $O_2$ ,  $O_3$ ,  $O_4$ ,  $O_5$ ,  $O_6$ ,  $O_6$ ,  $O_8$ ,  $O_8$ ,  $O_8$ ,  $O_8$ ,  $O_9$ ,

#### Results

The results obtained for all gases except  $CH_4$  are shown in Figures 1 through 8 which show the absorption coefficients as a function of wavelength. In each case the path length was 10 cm and the reference cell was filled with either  $N_2$  or Ar.  $CO_2$ , CO,  $CH_4$ , and  $O_2$  were measured at near atmospheric pressure. The  $H_2O$  pressure was 20 mm and the strongly absorbing gases NO,  $N_2O$ ,  $NH_3$ , and  $SO_2$  were measured at lower pressures. The pressures employed in each case are listed in Table I. For  $CH_4$  no

absorption could be detected, showing that the absorption coefficient is less than 10<sup>-4</sup> cm<sup>-1</sup> throughout the entire range investigated.

### Discussion

The results for  $CO_2$ ,  $CO_2$ , and  $O_2$  (see Figures 1, 2, and 3) are discussed in detail elsewhere (<u>Harteck, Reeves, and Thompson</u>, 1963). For  $CO_2$  it is found that absorption does occur in the wavelength region above 1750 corresponding to the spin-forbidden dissociation into ground-state products. CO exhibits absorption in this region only weakly in the narrow Cameron bands corresponding to the excitation to the  $a^3\Pi$  level. The absorption coefficient of  $O_2$  in the wavelength region between 2050 and 4000 A must be less than  $10^{-4}$  cm<sup>-1</sup> since no absorption was observed in this region.

 $H_20$  shows continuously decreasing absorption for wavelengths longer than 1850 A. The present results complement those of <u>Watanabe</u> et al. (1953) and extend them by over two orders of magnitude as shown in Figure 4. Similar remarks apply to the continuous absorption by  $N_20$  (Figure 5).

For NH<sub>3</sub>, NO, and SO<sub>2</sub> (Figures 6, 7, and 8) which show band structure rather than continuous absorption, the values obtained for the absorption coefficients will vary with the instrumental resolution employed. Nevertheless reasonable agreement was found with the results of <u>Watanabe et al.</u> (1953) for NH<sub>3</sub> and NO and with those of <u>Golomb</u>, <u>Watanabe</u>, and <u>Marmo</u> (1962) for SO<sub>2</sub>. The lack of absorption observed for CH<sub>4</sub> is in agreement with the results of <u>Watanabe et al.</u> (1953) since extrapolation of their data indicates that the absorption coefficient should be below 10<sup>-4</sup> cm<sup>-1</sup> at these wavelengths.

# References

- Golomb, D., K. Watanabe, and F. F. Marmo, Absorption Coefficients of Sulfur Dioxide in the Vacuum Ultraviolet, J. Chem. Phys. <u>36</u>, 958, 1962.
- Harteck, P., R. Reeves, Jr., and B. A. Thompsom, Photochemical Problems of the Venus Atmosphere, NASA TN D-1984, 1963.
- Watanabe, K., M. Zelikoff, and E. C. Y. Inn, Absorption Coefficients of Several Atmospheric Gases, <u>AFCRC Tech. Rpt.</u> 53-23, 1953.

Table I

Gas Pressures Employed for Absorption

# Measurements

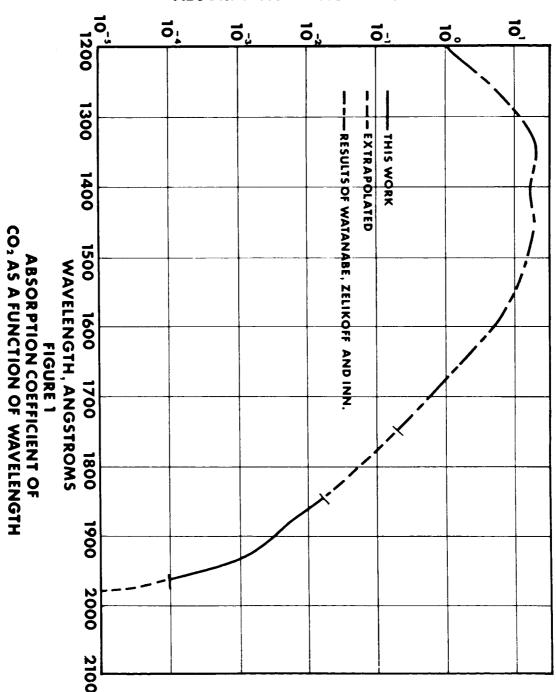
Gas	Pressure, mm.
C02	700
co	660
02	720
H <sub>2</sub> 0	20
N <sub>2</sub> 0	326, 2050 - 4000 A 6, 1850 - 2050 A
NH <sub>3</sub>	760, 2250 - 4000 A 100, 2250 - 4000 A 5, 1850 - 2150 A 0.1, 1850 - 2150 A
NO	<b>6*</b>
\$0 <sub>2</sub>	5, 2000 - 4000 A 0.259;1850 - 2200 A
CH4	750

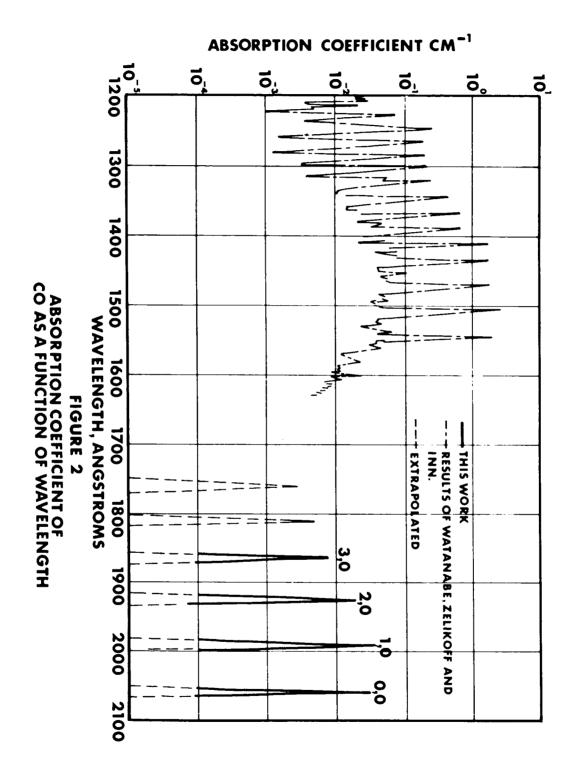
 $<sup>\</sup>ensuremath{^{\star}}\xspace \text{NO}$  was measured at low pressures to avoid interferences from dimerization.

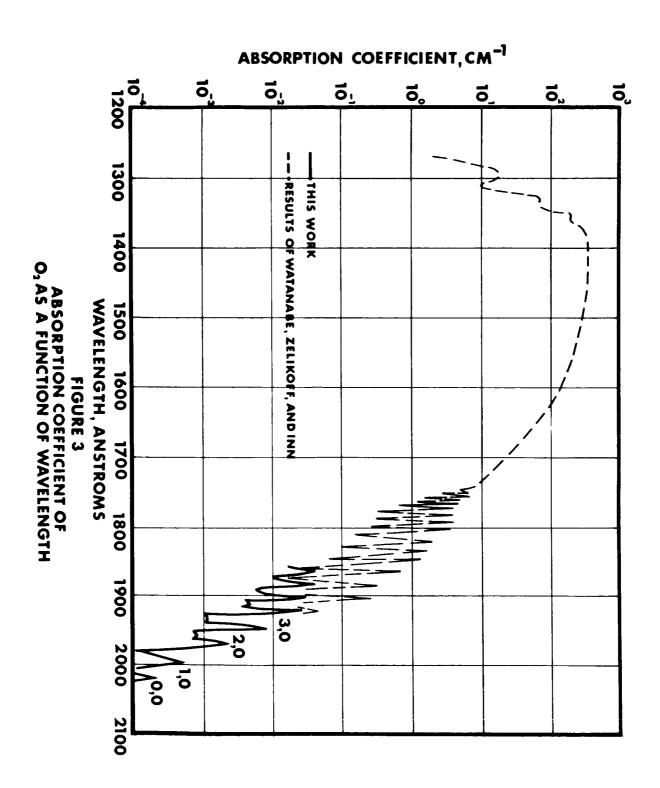
# List of Figures

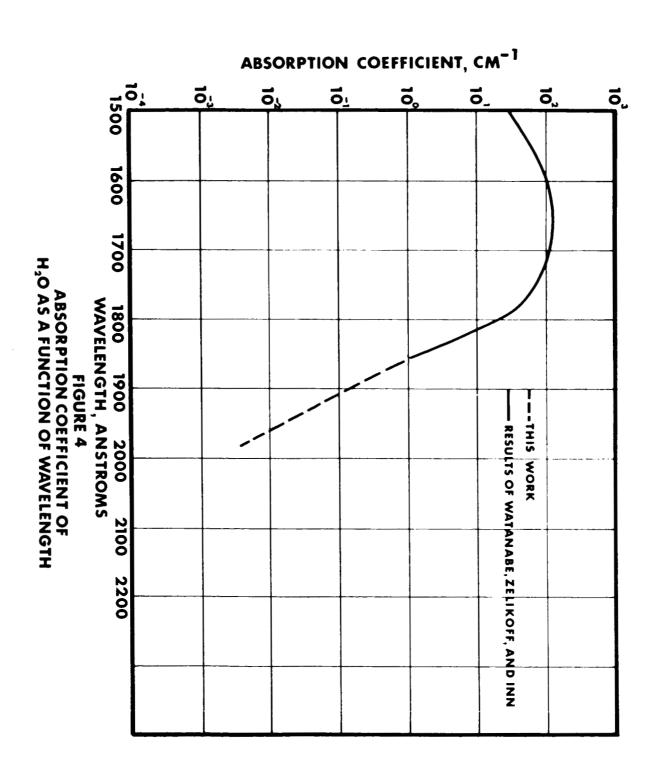
FIGURE	1	Absorption	Coefficient	of	CO <sub>2</sub> As a Function of Wavelength.
FIGURE	2	Absorption	Coefficient	of	CO As a Function of Wavelength.
FIGURE	3	Absorption	Coefficient	of	0 <sub>2</sub> As a Function of Wavelength.
FIBURE	4	Absorption	Coefficient	of	H <sub>2</sub> O As a Function of Wavelength.
FIGURE	5	Absorption	Coefficient	of	N <sub>2</sub> 0 As a Function of Wavelength.
FIGURE	6	Absorption	Coefficient	of	NH <sub>3</sub> As a Function of Wavelength.
FIGURE	7	Absorption	Coefficient	of	NO As a Function of Wavelength.
FIGURE	8	Abs <b>o</b> rption	Coefficient	of	SO <sub>2</sub> As a Function of Wavelength.

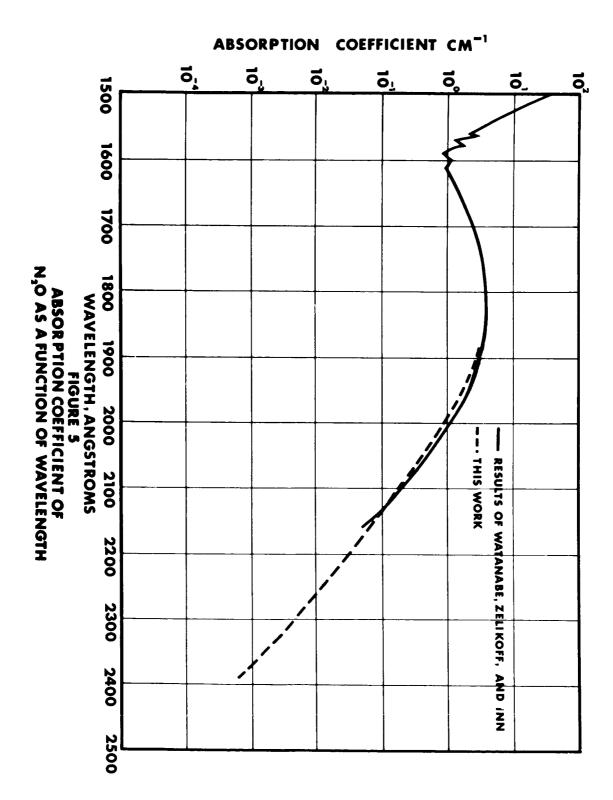
# ABSORPTION COFFICIENT CM-1

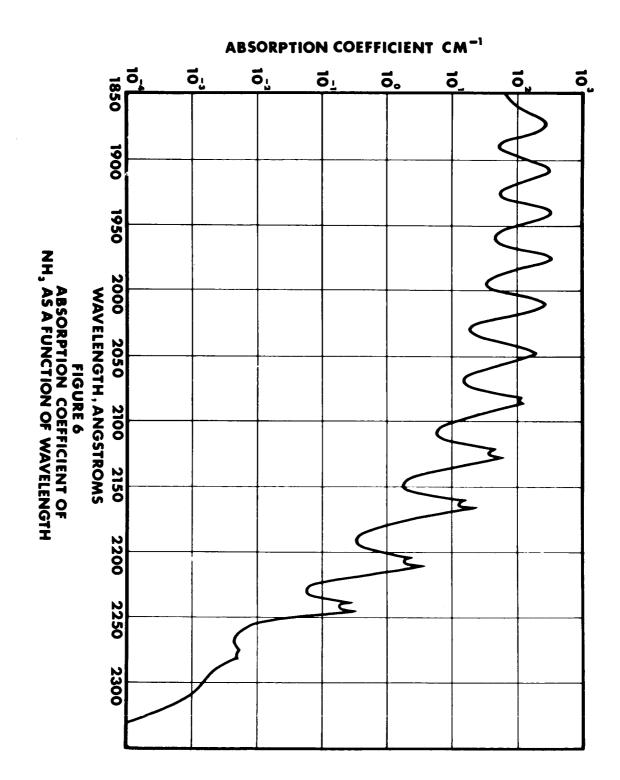




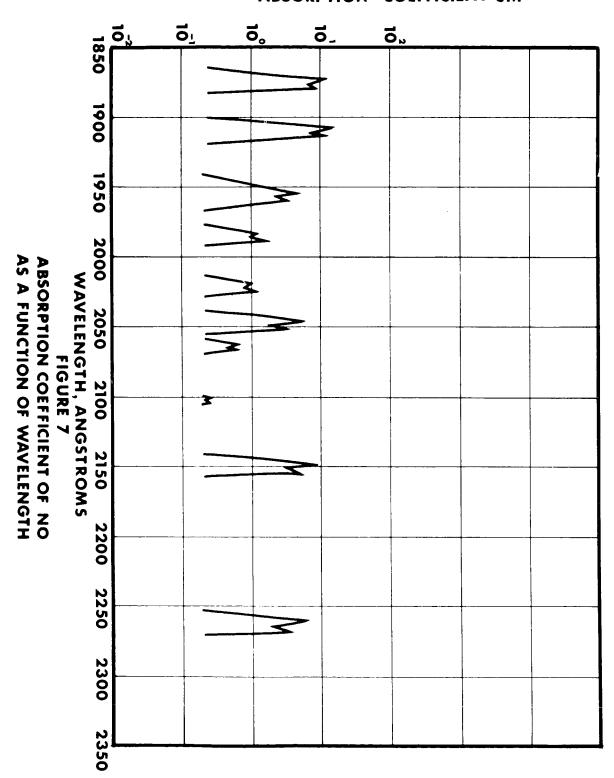




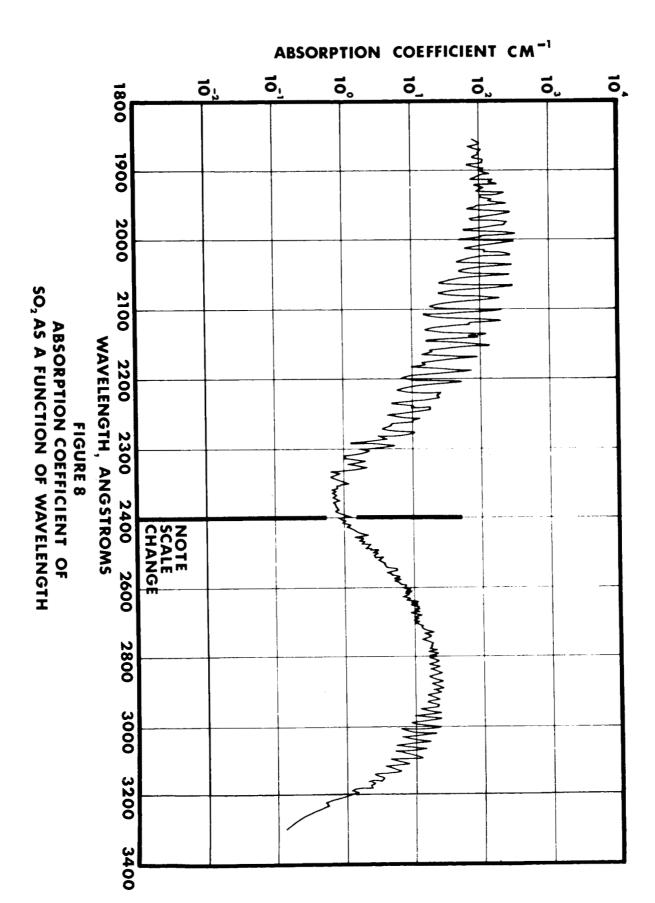




# ABSORPTION COEFFICIENT CM-1



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